

# Alpha decay of $^{114}\text{Ba}$

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In recent years, intense experimental and theoretical research has been devoted to investigate nuclei near the doubly magic nucleus  $^{100}\text{Sn}$  ( $Z=N=50$ ). In this region of the nuclear chart, an island of  $\alpha$  emission occurs, covering neutron-deficient isotopes of tellurium ( $Z=52$ ) through cesium ( $Z=55$ ). Alpha emission is a rich source of nuclear structure information [1]. The  $\alpha$ -particle energy  $E_\alpha$ , corrected for the recoil effect, yields the difference between the masses of mother and daughter nucleus. Above  $^{100}\text{Sn}$ , this quantity directly relates to the  $Z=N=50$  shell-strength. Moreover, the information on energy, half-life and  $\alpha$ -decay branching ratio ( $b_\alpha$ ) yields the reduced widths for this disintegration mode ( $W_\alpha$ ). Measurements of  $W_\alpha$  values may shed light on the question whether superallowed  $\alpha$  decay occurs for nuclei beyond  $^{100}\text{Sn}$ , with protons and neutrons occupying identical  $d_{5/2}$ ,  $d_{3/2}$  and  $g_{7/2}$  orbitals. Moreover, these nuclei, and in particular  $^{114}\text{Ba}$ , are predicted to be promising candidates for the observation of cluster ( $^{12}\text{C}$ ) radioactivity. The decay of  $^{114}\text{Ba}$  had already been investigated in previous experiments [2-4], but only lower limits for  $\alpha$  and  $^{12}\text{C}$  partial half-lives could be established [2]. In this contribution we report on preliminary results obtained in a reinvestigation of the  $\alpha$ -decay properties of  $^{114}\text{Ba}$ .

The experiment was performed at the GSI On-line Mass Separator.  $^{114}\text{Ba}$  was produced in fusion-evaporation reactions induced by a 284 MeV  $^{58}\text{Ni}$  beam on 2.0-3.8 mg/cm<sup>2</sup> thick  $^{58}\text{Ni}$  targets. The reaction products were stopped inside a high-temperature cavity ion source in two tantalum catcher foils, from which most recoils – at the ion source temperature of about 2400 K – are swiftly released as thermalized atoms. Chemical selectivity for barium isotopes was achieved by adding  $\text{CF}_4$  and thus using on-line fluorination. The ionization takes place selectively in the fluorination sideband of  $\text{BaF}$ , all contaminants including  $\text{CsF}$  being reduced to levels well below  $10^{-5}$  [5]. The  $\text{BaF}^+$  ions were accelerated to 55 keV, mass-separated and focused alternately onto two carbon foils, each one being viewed by a  $\Delta E$ -E silicon telescope.

The search for  $\alpha$  emission was performed using the  $\Delta E$  spectrum registered in anticoincidence with the E detector. Figure 1 shows the corresponding spectrum obtained from both telescopes during a measurement time of 55.6 hours. Three  $\alpha$  lines were observed at the energies  $3410\pm 40$ ,  $3740\pm 30$  and  $4160\pm 30$  keV, containing 38, 21 and 22 events, respectively. We assign the lowest-energy line to the  $\alpha$  decay of  $^{114}\text{Ba}$ . This assignment is based on the  $Z$  and  $A$  selectivity that was reached combining the fluorination ion source with mass-separation. The higher-energy members of this triplet are assigned to the known [6]  $\alpha$  lines from the decay of  $^{110}\text{Xe}$  and  $^{106}\text{Te}$ . Taking into account the recoil correction, we obtained a  $Q_\alpha$  value of  $3540\pm 40$  keV for  $^{114}\text{Ba}$ . This result, together with the known [6]  $Q_\alpha$  values of  $^{110}\text{Xe}$  and  $^{106}\text{Te}$ , yields a  $Q$  value of  $19000\pm 60$  keV

for  $^{12}\text{C}$  decay of  $^{114}\text{Ba}$ , which is important to obtain experimentally relevant predictions for this decay mode.

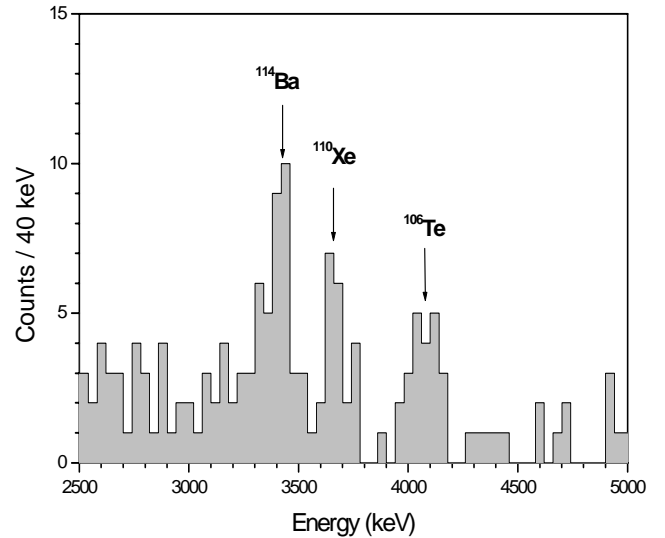


Figure 1. Section of the  $\Delta E$  spectrum from both telescopes, taken in anticoincidence with the related E detectors.

By comparing the 38  $\alpha$  events observed with the 980  $\beta$ -delayed protons ( $\beta p$ ) detected simultaneously in  $\Delta E$ -E coincidence, the  $b_\alpha$  value of  $^{114}\text{Ba}$  was determined to be  $(9\pm 3)\cdot 10^{-3}$ . This result takes into account the detection efficiencies of the telescopes and the known  $\beta p$  branching ratios for  $^{114}\text{Ba}$  [4] and its daughter  $^{114}\text{Cs}$  [7]. The new  $b_\alpha$  value is somewhat larger but still compatible with the previously obtained upper limit of  $3.7\cdot 10^{-3}$  (68% c.l.) [2]. Moreover, the analysis of time correlations will allow us to deduce the half-life of  $^{110}\text{Xe}$  and the  $b_\alpha$  values for  $^{110}\text{Xe}$  and  $^{106}\text{Te}$ .

## References

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